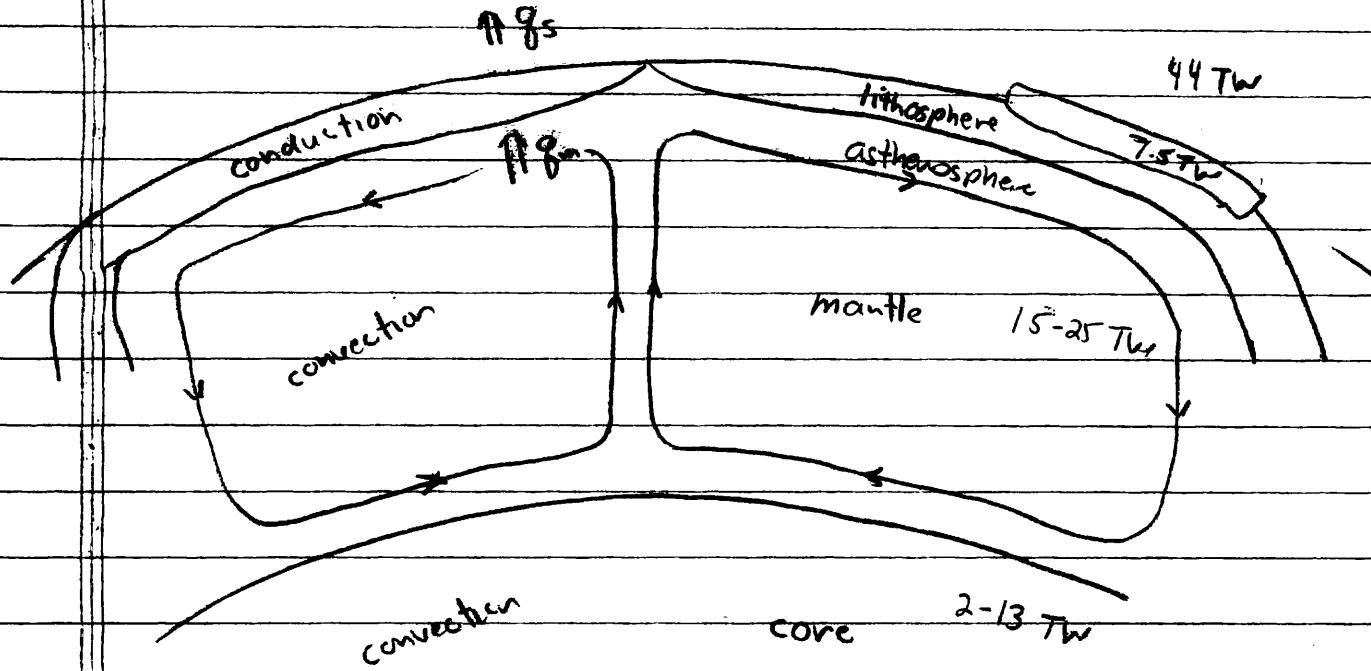


HEAT CONDUCTION

(Geodynamics, Chapter 4)



- decay of radioactive elements produces heat which must escape to the surface of the earth
- secular cooling also contributes to surface heat flux
- modes of heat transport:

$$\text{radiation} \quad q \propto T^4$$

$$\text{conduction} \quad q \propto \nabla T$$

$$\text{convection} \quad q \propto V \cdot \nabla T$$

- today's lecture

What is the average surface heat flow q_s ?

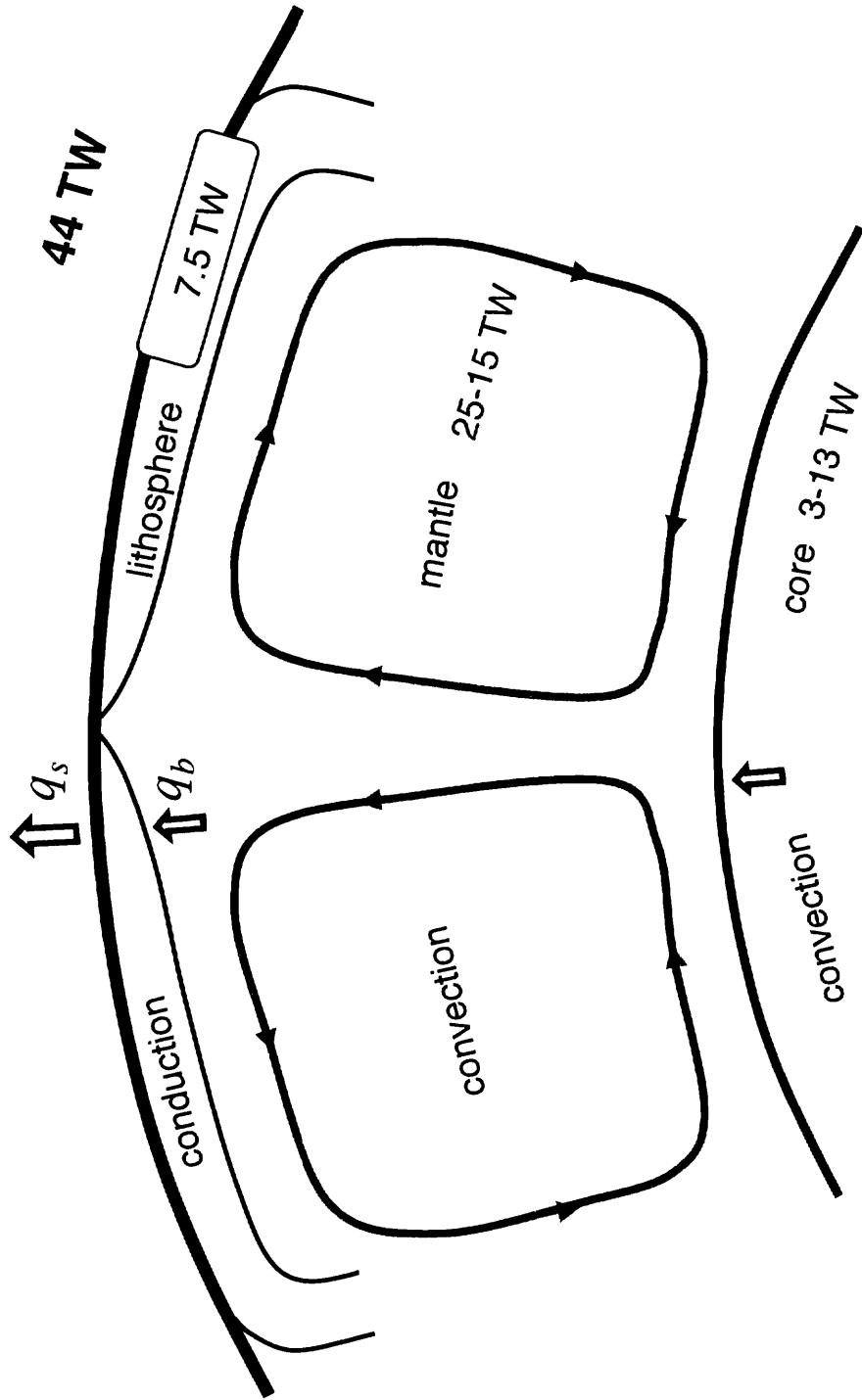
What is the average heat flow into the base of the lithosphere q_m ?

How does temperature increase with depth in continents and oceans?

How thick is the lithosphere and how does lithospheric thickness vary?

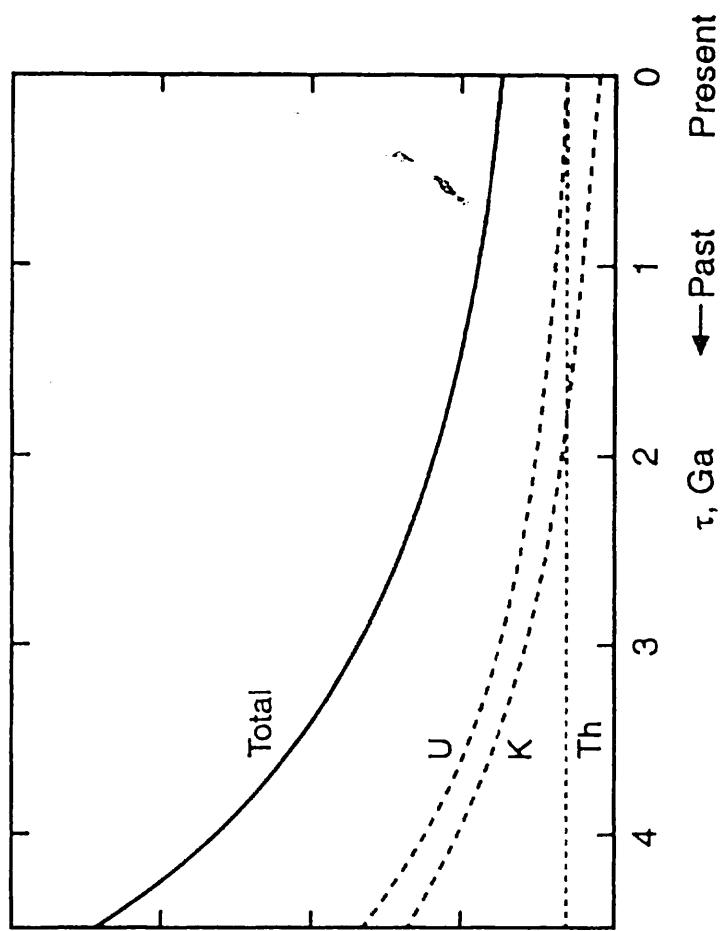
global heat budget

(1,2)



-5 HEAT GENERATION BY THE DECAY OF RADIOACTIVE ELEMENTS

137



4-4 Mean mantle heat production rates due to the decay of the radioactive isotopes of U, Th, and K as functions of time measured back from the present.

her para-
find that
er billion
ann man

plotted as a function of time before the present in Figure 4-4. The past contributions of the individual radioactive elements are also shown. We see that the rate of heat production 3×10^9 yr ago was about twice the

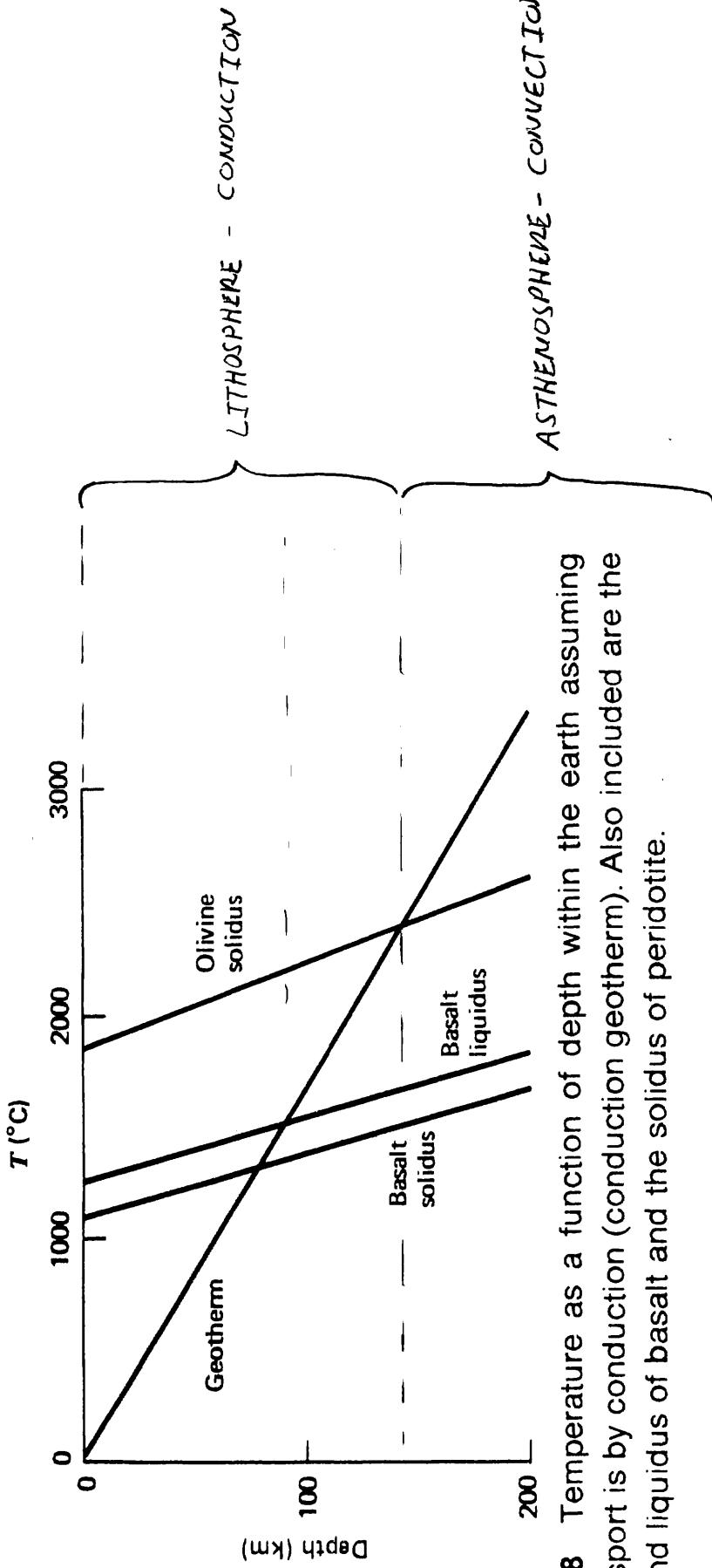


Figure 4-8 Temperature as a function of depth within the earth assuming heat transport is by conduction (conduction geotherm). Also included are the solidus and liquidus of basalt and the solidus of peridotite.

(3)

Conduction - Transfer of heat by molecular collisions. It is a diffusive process where heat flows from areas of higher T to lower T .

Convection - Transfer of heat by the motion of the rocks in the mantle.

Fourier's law of heat conduction

(see also: Carslaw, H. S. and J. C. Jaeger, Conduction of Heat in Solids, Oxford, 1959)

$$q = -k \frac{\partial T}{\partial z}$$

q - heat flow per area W/m^2

k - thermal conductivity W/m K
(material property)

T - temperature $^\circ\text{K}$ or $^\circ\text{C}$

For most solids, the thermal conductivity is relatively insensitive to temperature and it is nearly isotropic.

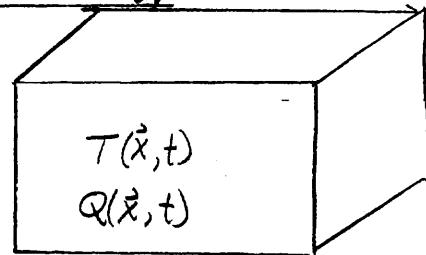
$$\vec{q}(\vec{x}) = -k \nabla T$$

heat flows in the direction of decreasing temperature.

Using the law of energy conservation one can derive the equation for diffusion of heat in an isotropic solid.

(4)

Heat Equation - Conservation of Energy



$$-\nabla \cdot k \nabla T + \rho C_p \vec{V} \cdot \nabla T + \rho C_p \frac{\partial T(\vec{x}, t)}{\partial t} = Q(\vec{x}, t)$$

divergence
heat flux advection of
heat change in
temp. with
time internal
heat generation

k - thermal conductivity $\text{W/m}^{\circ}\text{K}$

$Q(\vec{x}, t)$ - heat generation per volume W/m^3

ρ - density kg/m^3

C_p - heat capacity at constant P $\text{J/kg}^{\circ}\text{K}$

$\chi = \frac{k}{\rho C_p}$ - thermal diffusivity $\frac{\text{m}^2}{\text{s}}$

$H = Q/\rho$ - heat generation per kg $\text{W/kg}^{\circ}\text{K}$

divide through by ρC_p

$$\vec{V} \cdot \nabla T + \frac{\partial T}{\partial t} = \chi \nabla^2 T + \frac{Q}{\rho C_p}$$

The thermal diffusivity of a solid is a measure of the amount of time it takes for a heat pulse to travel a distance l . $\tau = \frac{l^2}{\chi}$

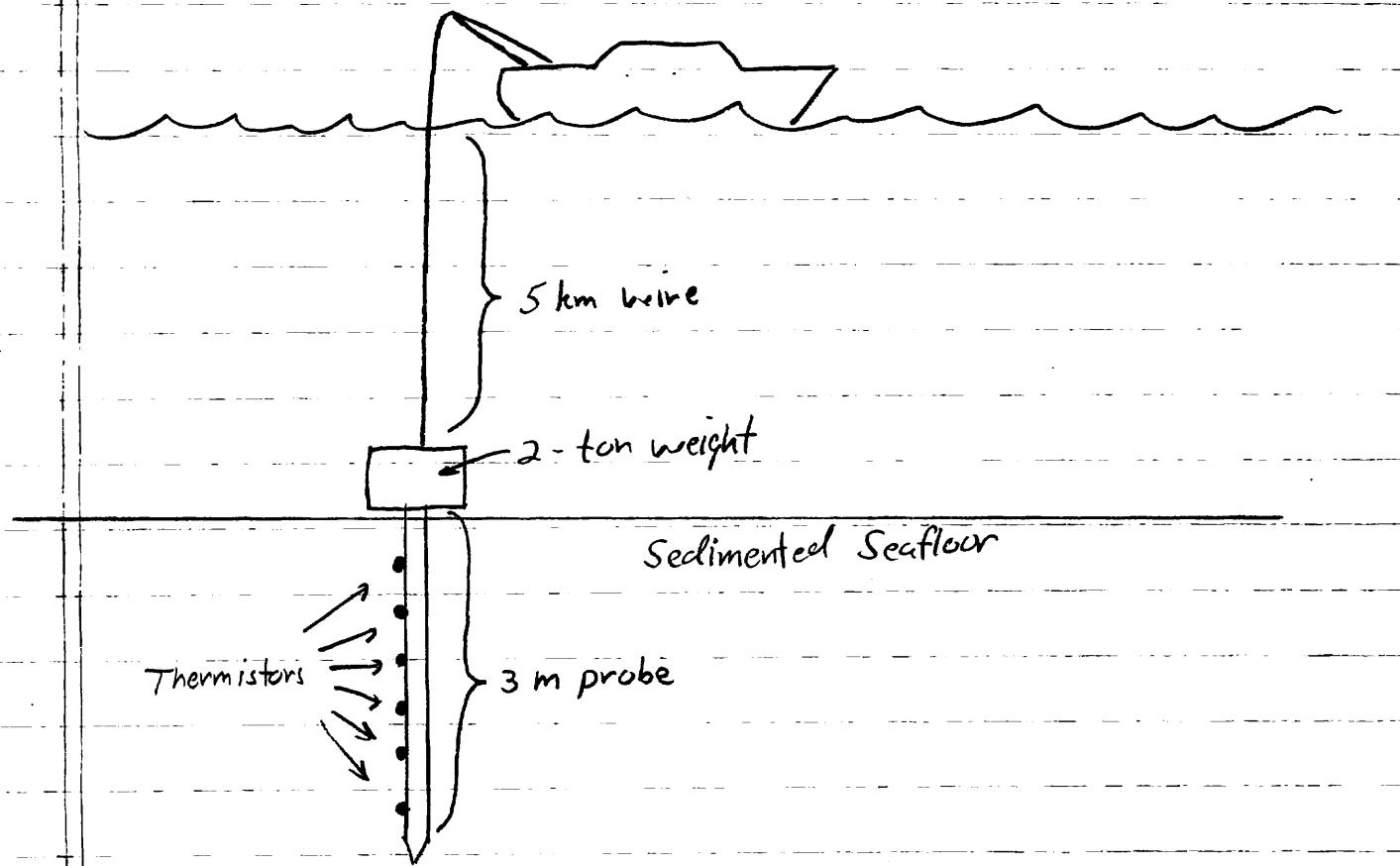
(5)

Measurement of Heat Flow from the Earth:

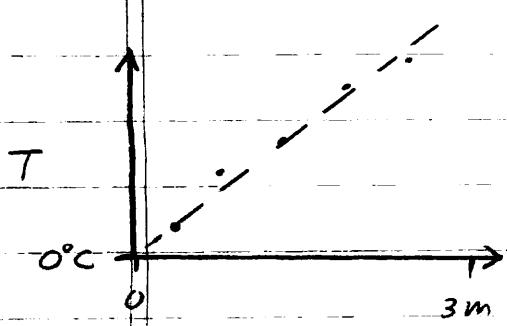
$$q_s = -k \frac{\partial T}{\partial z}$$

Need to measure surface temperature gradient and the surface thermal conductivity.

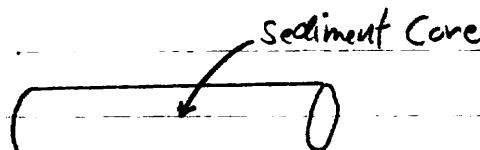
Ocean Measurements:



Temperature Gradient



Conductivity



measure conductivity in the lab.

SCLATER ET AL.: OCEANIC AND CONTINENTAL HEAT FLOW

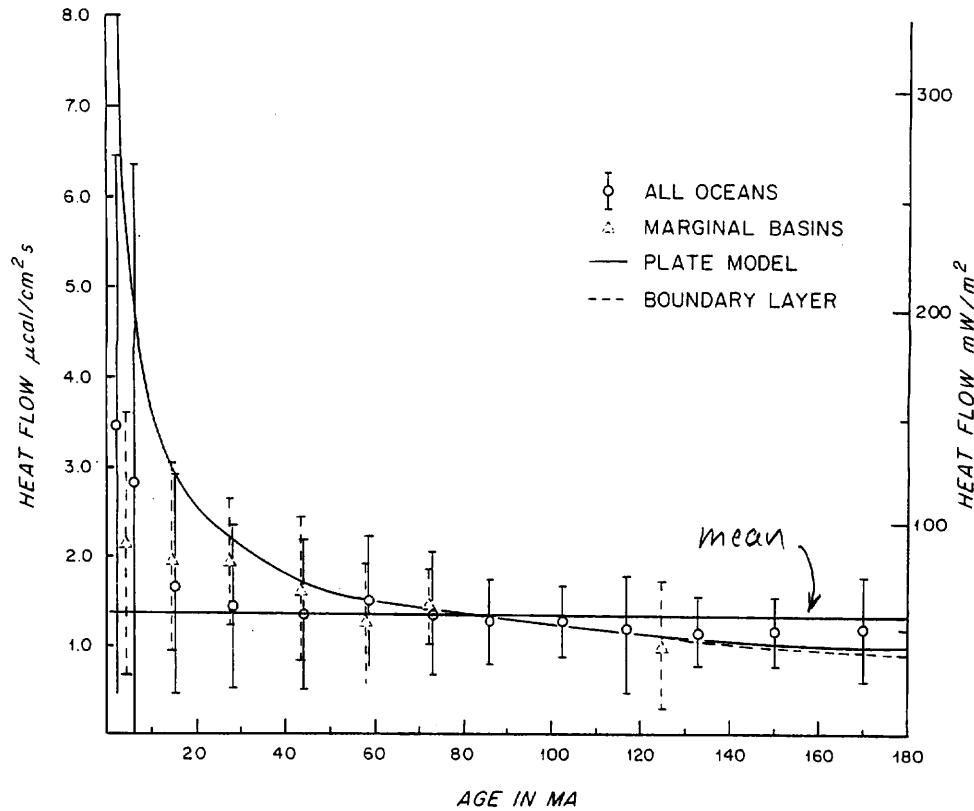
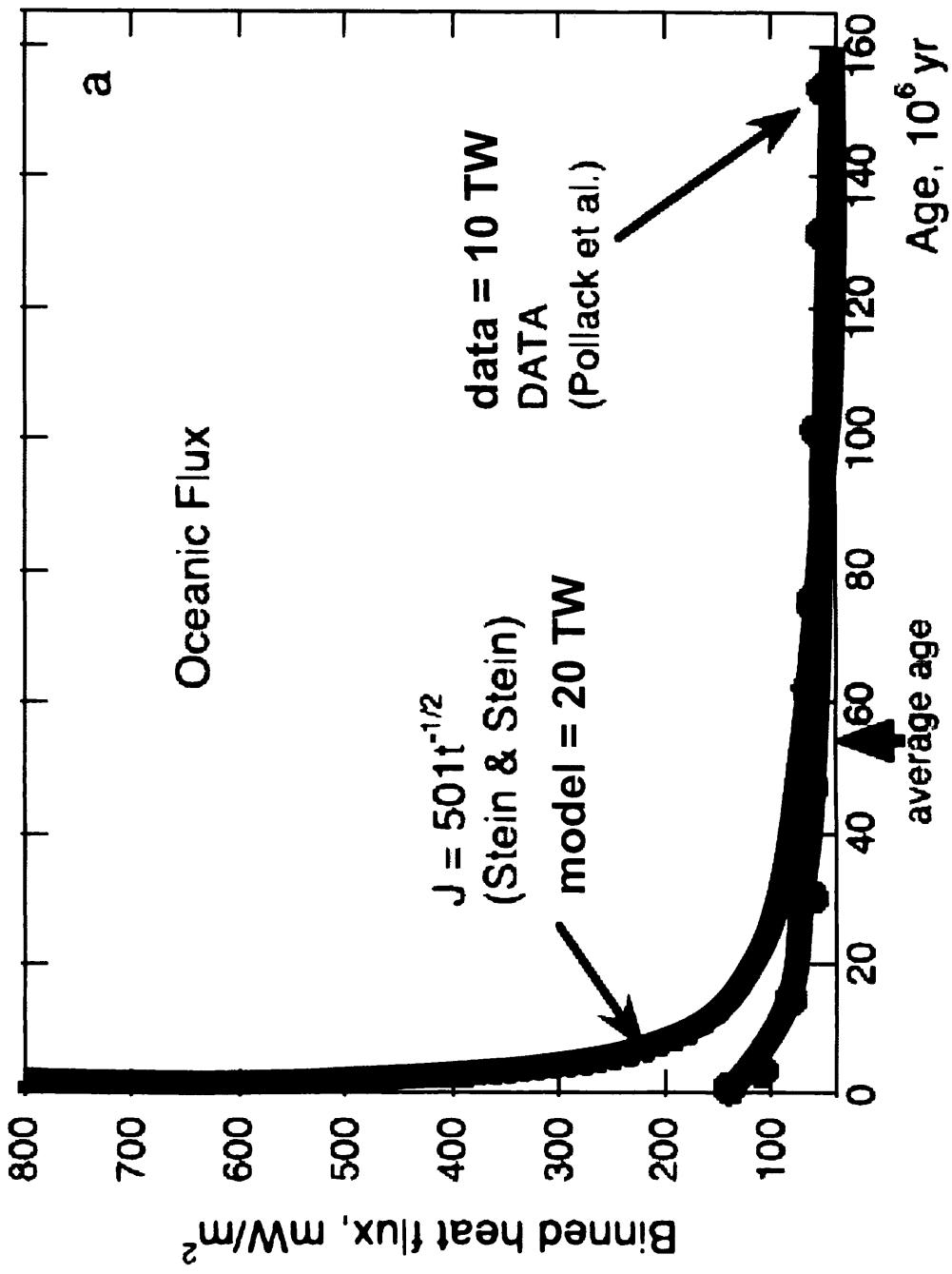


Fig. 4a. Mean heat flow and standard deviation for all the oceans and for the marginal basins as a function of age. Also shown are the expected heat flows from the plate and boundary layer models.

- Mean oceanic heat flow is $\sim 66 \text{ mW/m}^2$
- Heat flow decreases with increasing seafloor age.
- For ages $< 20 \text{ Ma}$ the oceanic conductive heat flow is than expected!

$$\text{heat flow vs age} \rightarrow q(t) = k \frac{\partial T}{\partial z} \rightarrow q(t) \cong 480 t^{-1/2}$$

A.M. Hofmeister, R.E. Criss / Tectonophysics 395 (2005) 159–177



(7)

Continent Measurements:

Seasonal Temperature Variations

Glacial Cycle Temperature Variations

$$\text{measure } \frac{\partial T}{\partial z} \rightarrow \{$$

> 300 m deep

Measure conductivity of core in the lab: 

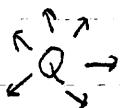
* on board

Continental crust rocks contain higher concentrations of radioactive elements.

$$\dot{g}_s = \dot{g}_c + \dot{g}_m$$

$$\uparrow \dot{g}_s$$

Crust
(~30 km)



Lithosphere

Mantle

$$\uparrow \dot{g}_m$$

Asthenosphere

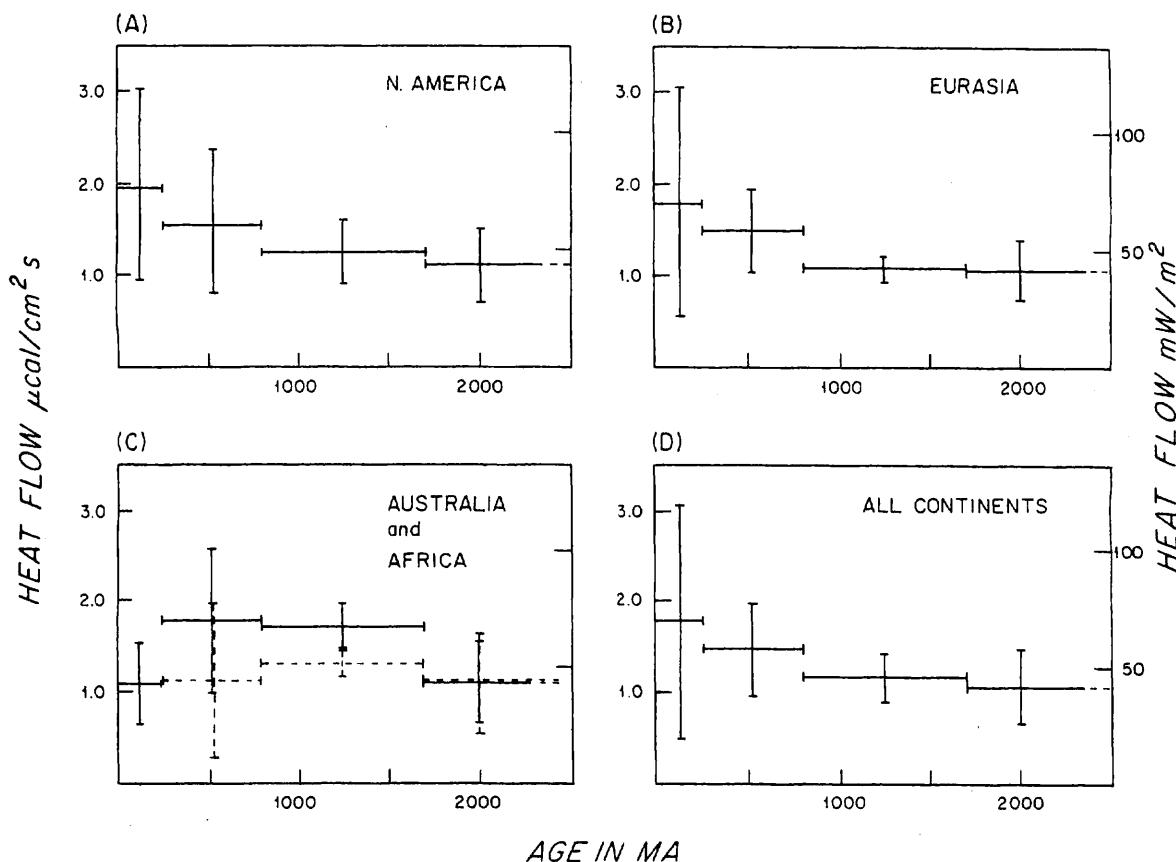


Fig. 16. Heat flow as a function of age for (a) North America, (b) Eurasia, (c) Australia and Africa (dashed lines), and (d) all continents.

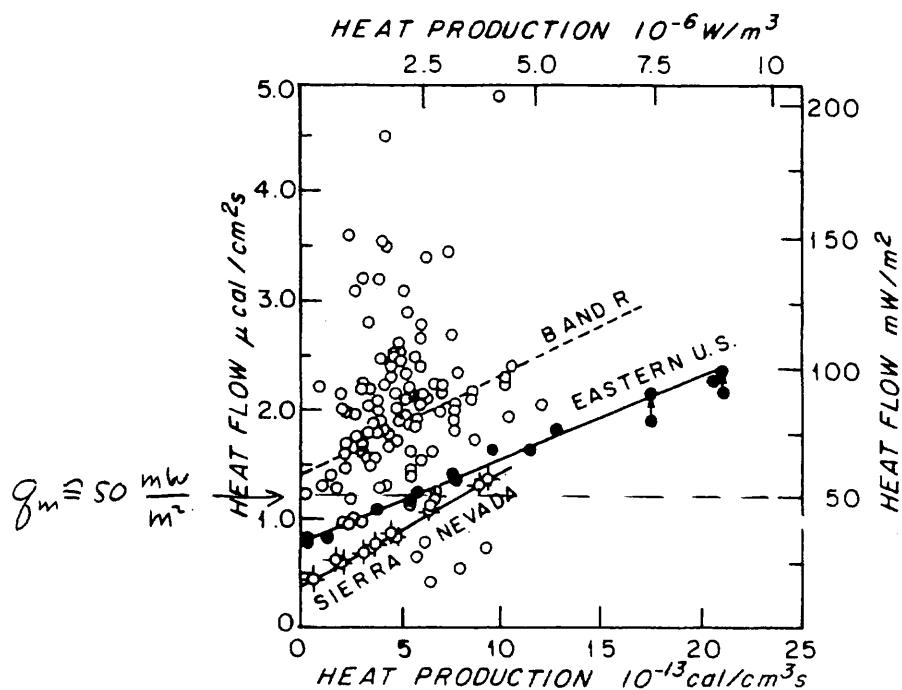


Fig. 17. Observations of heat flow and radioactive heat production from crystalline rocks in the United States; linear regression curves are from Roy et al. [1968] for the Basin and Range (dashed curve), eastern United States, and Sierra Nevada provinces. Solid circles represent points east of the Great Plains, and crossed circles represent points interior to the Sierra Nevada physiographic province.

(9)

Let's assume that the heat production of crustal rocks decreases exponentially with depth.

$$Q(z) = Q_s e^{-z/h_r}$$

↑ ↑ →
 heat generation measured assumed decay
 at depth surface heat with depth
 generation

$$\nabla^2 T = -\frac{Q}{k}$$

What is the crustal contribution to the continental heat flow?
 What does a typical continental geotherm look like?

Assume steady heat conduction.

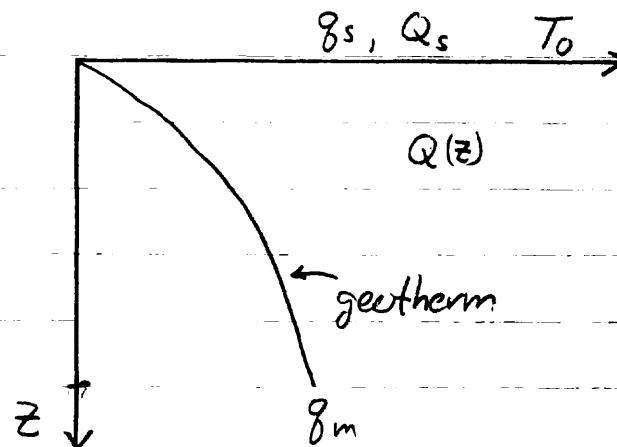
$$\frac{\partial^2 T}{\partial z^2} = -\frac{Q_s}{k} e^{-z/h_r}$$

$$\left. \frac{\partial T}{\partial z} \right|_0 = -\frac{Q_s}{k}$$

$$\left. \frac{\partial T}{\partial z} \right|_\infty = -\frac{Q_s}{k} \quad \int_0^z e^{-s/h_r} ds = -h_r [e^{-\frac{z}{h_r}} - 1] = h_r [1 - e^{-\frac{z}{h_r}}]$$

$$\text{Integrate Once: } -k \frac{\partial T}{\partial z} = g(z) = -Q_s h_r [1 - e^{-\frac{z}{h_r}}] + C_1$$

$$\text{Use boundary condition } g(\infty) = g_m$$



(10)

$$\text{Heat Flow: } q(z) = Q_s h_r e^{-\frac{z}{h_r}} + q_m$$

Surface Heat Flow: $q_s = Q_s h_r + q_m$

↑ ↑ ↑
 surface heat surface heat mantle heat
 flow production flow
 X
 effective
 thickness

Show Heat flow vs Heat Production Again

Show Global Heat Flow

h_r = slope of q_s vs Q_s plot $\sim 10 \text{ km}$

$q_m = q_s - Q_s h_r$ - reduced heat flow

Theoretical Continental Geotherm:

Integrate again and use $T(0) = T_0$ boundary condition

$$T(z) = T_0 + \frac{q_m z}{k} + \frac{Q_s h_r^2}{k} [1 - e^{-\frac{z}{h_r}}]$$

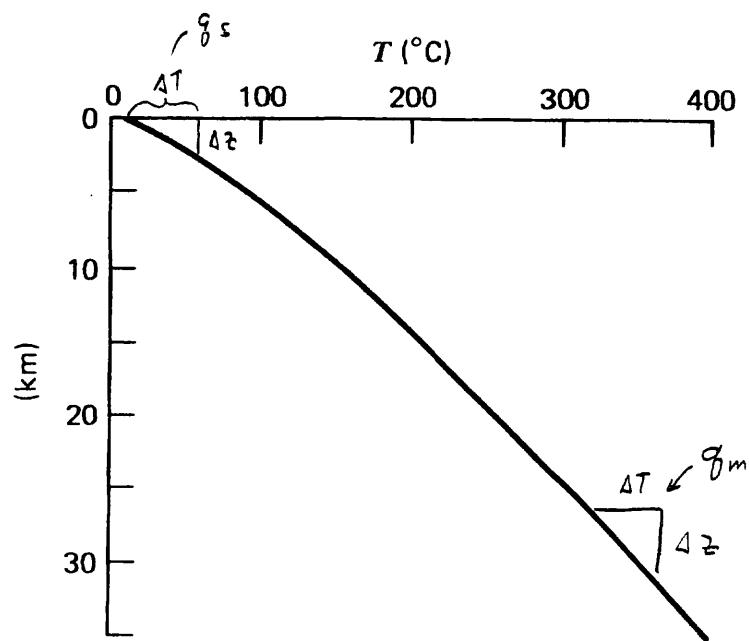


Figure 4-12 A typical geotherm in the continental crust.

TABLE 12. Summary of Area and Heat Loss Information

	Area, 10^6 km^2	Heat Loss, $10^{10} \text{ cal/s} (10^{12} \text{ W})$
Continents (including effect of volcanoes)	149.3	208 (8.8)
Continental shelves	52.2	67 (2.8)
Total (continent plus shelves)	201.5	275 (11.5)
Deep oceans	281.7	656 (27.4)
Marginal basins	26.9	71 (3.0)
Total (oceans plus marginal basins)	308.6	727 (30.4)
Worldwide values	510.1	1002 (42.0) ← TOTAL LOSS
Heat loss by water flow and conduction through bare rock		241 (10.1)
Heat loss by plate creation not including radiogenic heat loss		626 (26.2)
Mean heat flow		
Continents and shelves		1.37* (57)
All oceans (observed)		1.57* (66)
All oceans (theoretical)		2.36* (99)

Average Heat Flow

*Heat flow in $10^{-6} \text{ cal/cm}^2 \text{ s}$ (mW/m^2).

Heat Flow Considerations

(from Jerry Schubert's lecture at IGPP, Fall, 2000)

- Heat flow from the Earth is 44 TW.
- Heat flow from radiogenic heat production in the continents is 7.5 TW.

This is uncertain since we do not really know the amount of U, Th, and K in the continents, nor do we know the heat flow from the mantle into the base of the continental lithosphere.
- Heat flow from the mantle is 36.5 TW =
Heat from the core +
Heat from mantle cooling +
Heat from radioactivity.

Heat Flow From the Mantle 36.5 TW

(from Jerry Schubert's lecture at IGPP, Fall, 2000)

Heat flow from core: 2.4 TW [Sleep, Davies]

13.4 TW [Malamud & Turcotte]

Urey Number = $\frac{\text{total heat}}{\text{mantle cooling}}$

Guy Masters
says the
mantle needs to
run the
diverso

Heat flow from mantle cooling

if $U_r = 0.75$ 9.1 TW

Heat produced by radiogenic sources in the mantle is 14 - 25 TW?

Heat Flow Considerations

(From Jerry Schubert's lecture at IGPP, Fall, 2000)

- Kellogg et al., [1999] have argued that the MORB source region is so depleted in radioactivity (heat production) that were it identical to the entire mantle, it could not supply the heat observed to be escaping through the earth's surface.
- According to this argument there must be an isolated region of the mantle with a higher concentration of Ur, Th, and K compared to the MORB source region.
- Kellogg et al., [1999] hypothesize that this region is a compositionally distinct (undepleted, rich in radiogenic elements, ${}^3\text{He}$ and ${}^{40}\text{Ar}$), variable thickness, separately convecting layer buried at the bottom of the lower mantle. However, the constraints on mantle radioactivity are not so certain.

