

Experiment of Ionospheric Corrections to ALOS L-Band Interferograms

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Outline

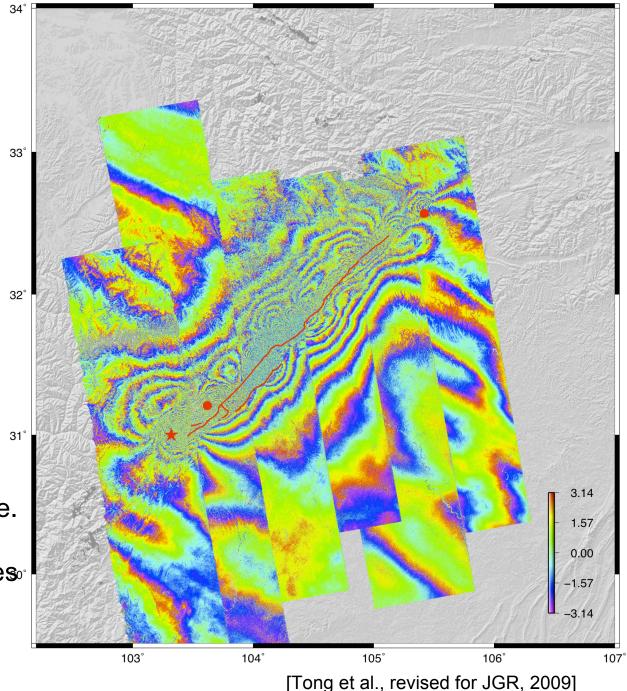
- Example: Interferograms of 2008 Wenchuan earthquake displayed waves and ramps, which is probably associated with ionospheric disturbance.
- Attempt to correct the ionospheric phase for L-band ALOS PALSAR data using available ionosphere models.
 - 20 images near PFO, Southern California
 - Theory and method
 - Ionospheric models: GIM, USTEC, TEC-DAWN
- Results and a summary

Example

 Ionospheric phase ramps and waves are evident in coseismic interferometry of 2008 Wenchuan earthquake.

• Ionospheric waves cause azimuth shifts ^{3*} resulting in wave-like areas of lower coherence.

• Similar phase anomalies are observed in other areas: Japan, Canada, Antarctic.



ALOS L-band PALSAR Data

- T213 F650~670
- 20 ascending acquisitions
- Time span: 2006 to 2009



- Process 11 interferograms to residual phase, many of them exhibit phase ramps and waves.
- Select short spatial baseline pairs (< 514m) to reduce potential orbital error.
- Neglect the interferograms that have large turbulent tropospheric signature.
- Motivation: Can global or regional ionospheric models be used to correct the phase ramps ?

Theory : Effects of Ionosphere on Range

index of refraction

$$n = \sqrt{\varepsilon} = \sqrt{1 - \frac{\lambda^2 e^2 N_e}{4\pi^2 m \varepsilon_0 c^2}} \approx 1 - \frac{1}{2} \frac{\lambda^2 e^2 N_e}{4\pi^2 m \varepsilon_0 c^2} = 1 - K \lambda^2 N_e$$

phase velocity > *c*

 $v_p = c / n$

vertical travel time change

$$\Delta \tau = \int_{0}^{H} (1/v_{p} - 1/c) dz = -\frac{K\lambda^{2}}{c} \int_{0}^{H} N_{e}(z) dz = -\frac{K\lambda^{2}}{c} TEC$$

range change (1-way vetical) $\Delta \rho = -K\lambda^2 TEC$

L-band example (1-way vertical)

$$\Delta \rho = -25 cm * TECU$$

$$N_{a}$$
 - electron density

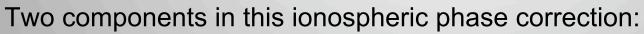
- e^{e} electron charge
- *m* electron mass
- λ radar wavelength
- *c* speed of light
- ε_o permittivity of free space

[Gray et al., 2000; Rees, 2001]

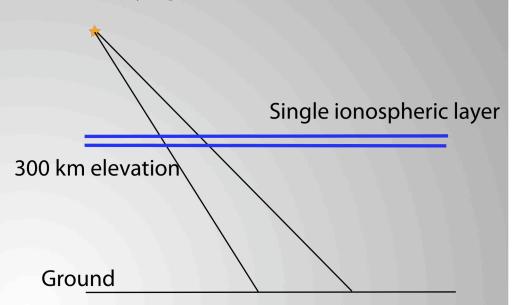
Method

• Extract vertical TEC value from three TEC models respectively.

- Compute geographic coordinates at ~300 km elevation that correspond to the ground-level image based on looking angles and the satellite orbit.
- Interpolate the TEC model and produce ionospheric phase corrections based on the formula.

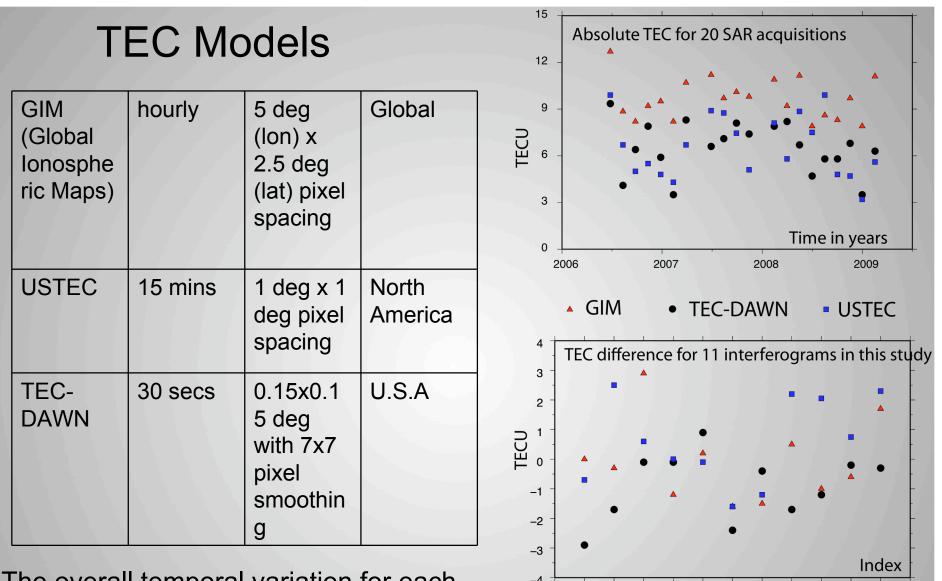


- A. Spatial variations in TEC **difference** will map directly into range differences.
- B. Uniform TEC **difference** across the area will map indirectly according to the increasing range across the swath.



 $\Delta \phi = -4\pi K \lambda * TEC_{vertical} / \cos(E)$

satellite is flying at ~ 700 km elevation



- The overall temporal variation for each model (RMS about the mean): 1.1~1.4 TECU
- The overall disagreement (RMS after differencing between models): 1.5~1.7 TECU

0

2

10

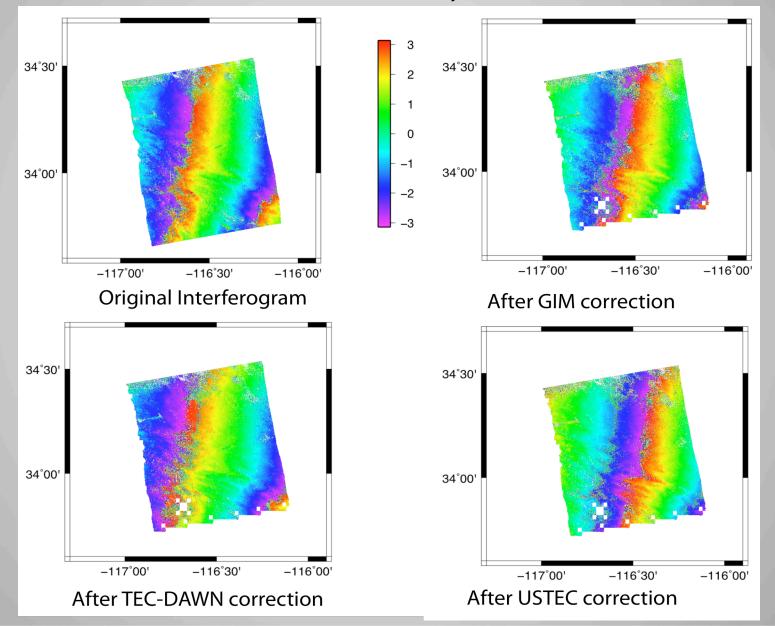
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11 12

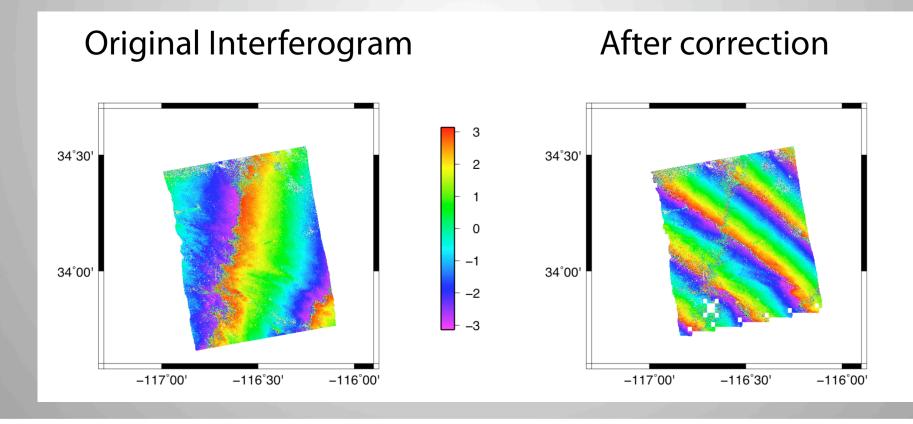
• GIM has an mean offset of 3.1 TECU compared to the other two models.

Results (Example with only a constant TEC difference)



Results (Example with a linear TEC model in both range and azimuth)

Fit a bilinear function based on finer resolution TEC-DAWN model to account for the spatial variation of TEC within one interferogram.
TEC-DAWN become more noisy near coastline which can be a problem.



Summary

- Typical ionospheric variations will produce ramps in phase across interferograms of 0~2 fringes that could be confused with orbit error.
- We were not successful in removing phase ramps in range using uniform TEC differences derived from three different ionospheric models.
- The three models provide a constant TEC difference but they have an overall disagreement at 1.5~1.7 TECU, which maps into 0.5~0.6 fringes in range.
- We also tried to correct range and azimuth ramps using the finer resolution TEC-DAWN model but were not successful.
- We conclude that none of the current TEC models are accurate enough to correct phase ramps in interferograms.